The Riemann-Hilbert method: from Toeplitz operators to black holes

Maria Cristina Câmara

CAMGSD-Instituto Superior Técnico

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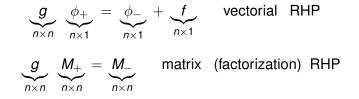
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$$\phi_{+}(t) + \phi_{-}(t) = f(t), \ t \in \mathbb{T}$$

$$\Leftrightarrow -\phi_{+} = \phi_{-} - f, \text{ on } \mathbb{T}$$

$$g \phi_{+} = \phi_{-} - f, \text{ on } \mathbb{T}$$

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Bounded Wiener-Hopf (or Birkhoff) factorization:

 $g = M_- M_+^{-1}$ $M_+^{\pm 1}$ analytic and bounded in \mathbb{D} $M_-^{\pm 1}$ analytic and bounded in $\mathbb{C} \setminus clos \mathbb{D}$

RH approach: to reduce a problem to the reconstruction of a function analytic in $\mathbb{C} \setminus \Gamma$ from jump conditions across Γ .

Applications in

- Diffraction problems
- Elastodynamics
- Singular integral equations
- Combinatorial probability
- Random matrices
- Orthogonal polynomials
- Integrable systems

Deift, Its, Kapaev, Novokshenov, Fokas, Ablowitz, Bleher, Östensson

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The Riemann-Hilbert method: a Swiss Army knife Marko Bertola (Concordia University), 2012

There are **no general methods to solve matrix RHP**. One has to develop custom-made methods, case by case.

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Considerable progress has been made in **explicit** factorisation methods.

M. C. Câmara, A. F. dos Santos and P. F. dos Santos: *Matrix Riemann-Hilbert* problems and factorization on Riemann surfaces, J. Funct. Anal. (2008).

M. C. Câmara, C. Diogo and L. Rodman: *Fredholmness of Toeplitz operators and corona Problems*, J. Funct. Anal. (2010).

M. C. Câmara, C. Diogo, Yu. Karlovich and I.M. Spitkovsky: *Factorizations, Riemann-Hilbert problems and the corona theorem*, J. London Math. Soc. 86 (2012).

M. C. Câmara, C. Diogo and I.M. Spitkovsky: *Toeplitz operators of finite interval type and the table method*, J. Math. Anal. Appl. (2014).

Spectral properties and kernels of Toeplitz operators (TO)

In the context of $L^2(\mathbb{T})$ or $L^2(\mathbb{R})$:

$$L^{2} = H_{+}^{2} \oplus H_{2}^{-} \qquad \qquad H_{2}^{\pm} = \mathcal{F}L^{2}(\mathbb{R}^{\pm})$$
$$P^{+}: L^{2} \longrightarrow H_{2}^{+}$$

Toeplitz operator:

 $\begin{array}{ll} T_g: H_2^+ \longrightarrow H_2^+ & g \in L^{\infty} \\ T_g \; \varphi_+ = P^+ g \, \varphi_+ & g \text{ is the symbol of the TO (it can be matricial).} \end{array}$

In $L^2(\mathbb{R})$, TO are unitarily equivalent, via the Fourier transform, to convolution operators on the half-line \mathbb{R}^+ .

Toeplitz operators are intimately related to RHP:

 Fredholmness, invertibility, the dimension of the kernel and the cokernel (and therefore their spectral properties) are determined by a RH factorisation of their symbol.

In particular: T_q is invertible $\iff g = M_- M_{\perp}^{-1}$

Progress in developing methods to explicitly solve RH factorisation problems goes hand in hand with progress in the spectral theory of Toeplitz operators



A Kernels of TO:

many important spaces of functions, such as model **spaces**, can be described as Toeplitz kernels. Toeplitz kernels consist of the solutions to a vectorial RHP

$$g\phi_+ = \phi_-$$

Some recent results taking this RH approach to Toeplitz kernels:

- new (and surprising) properties of all Toeplitz kernels

M. C. Câmara and J. R. Partington, *Near invariance and kernels of Toeplitz operators*, J. Anal. Math. (2014).

- generalisation of Hitt's and Hayashi's results on nearly invariant subspaces to a Banach space setting

M.C. Câmara and J.R. Partington, *Finite-dimensional Toeplitz kernels and nearly-invariant subspaces*, J. Operator Theory (2016).

- characterisation of the multipliers between Toeplitz kernels M.C. Câmara and J.R. Partington, *Multipliers between Toeplitz kernels* (2017).

Constructing new solutions to Einstein's field equations

- Einstein's field equations are **nonlinear** PDE's.
- They are very difficult to solve in general, so one must concentrate on special classes of solutions which exhibit symmetries.
- Reduced (2 dimensional) field equations:

$$d(\rho * A) = 0.$$

Here * denotes the Hodge dual (* $d\rho = dv$, * $dv = -d\rho$) and *A* is a matrix one-form $A = M^{-1}dM$ where $M(\rho, v)$ determines the solution to Einstein's equations.

 Following Breitenlohner-Maison's approach (1987) one can construct a linear system depending on an additional variable called the spectral parameter *τ*:

$$\tau(dX+AX)=*dX.$$

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$$d(\rho * A) = 0 \quad A = M^{-1} dM \tag{3}$$

$$\tau(dX + AX) = *dX \quad \text{Lax pair} \tag{4}$$

(3) is a compatibility condition for (4) (**integrable system**) provided that τ satisfies a certain differential equation

$$\implies \tau \omega = \tau v + \frac{\rho}{2}(1 - \tau^2)$$
 spectral curve

The **RH approach** ([CCMN]): to each solution of (3) we can associate a matrix $\mathcal{M}(\omega)$ such that

$$\mathcal{M}(\omega)|_{\omega = v + \frac{\rho}{2} \frac{1 - \tau^2}{\tau}} = M_{-}(\tau)M_{+}^{-1}(\tau) \qquad |\tau| = 1$$

$$M_+^{-1} = X, \quad \lim_{\tau \to \infty} M_-(\tau) = M(\rho, \nu)$$

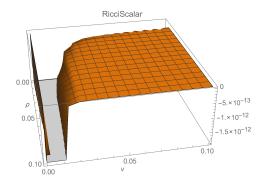
and conversely.

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Using explicit factorisation methods:

- we solve the long-standing problem of constructing extremal black holes within the RH formulation;
- by acting with elements of the Geroch group on the monodromy matrix of a "seed solution", we obtain **new solutions** to Einstein's equations.

A concrete example: Deforming the monodromy matrix corresponding to the solution that describes the near horizon region of an extremal black hole, we obtain a new solution. It is explicit, completely regular, and exhibits unexpected properties that one would not guess a priori.



[CCMN] M.C. Câmara, G.L. Cardoso, T. Mohaupt, S. Nampuri: *A Riemann-Hilbert approach to rotating attractors*, JHEP (2017).